



FIRST LIGHT IMAGING Green House Gas detection & observation from space in the SWIR band

Jean-Luc Gach – <u>jeanluc.gach@first-light.fr</u> Feb 2024

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Gas detection principle



Radiation Transmitted by the Atmosphere

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70

70

Hyperspectral Earth observation



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J band GHG spectral detection



High spectral resolution permits species disentenglement Absorption depth gives concentration High spatial resolution enhances detectivity (less dilution)

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Optical considerations

Diffraction limit





Diffraction limit

Theoretical spatial resolution limit for a 25cm telescope at variuous LEO altitudes (400 to 800km)



Necessary telescope diameter (cm) for a 10m ground resolution for varoius LEO altitudes (400 to 800km)

Acquisition speed impact on ground resolution

Ground resolution due to satellite displacement (LEO orbit) vs framerate

In the case of an hyperspectral imager, the satellite displacement x acquisition speed gives the Y ground resolution



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Conclusions

- Hyperspectral imaging is more favorable in the SWIR for high spatial resolution (among some other advantages)
- Longer wavelengths lead to high size and costly optics, high weight and additional launch costs. Not compatible with cubesats.
- Acquisition speed brings also a spatial resolution limit and a high framerate (>500 FPS) is mandatory for ~10m ground resolution and higher

What First Light Imaging can do

C-RED New Space



For more info visit: <u>www.first-light-imaging.com</u>

- SWIR sensitivity from 0.9 to 1.7 (or 2.2 μm)
 - Actual theoretical ground resolution of 6m with a 25cm telescope
 - Sensitivity to the 1.6um CO₂ and CH₄ absorption bands

Acquisition speed of 600FPS full frame

- Higher speed if cropped (less spectral bands), up to KHz framerates
- Not limiting for pushbroom hyperspectral imager applications
- Low noise, low cost, customizable and many other features !

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