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# Aeolus – 1<sup>st</sup> wind lidar in space

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Outline





# Objectives



# **Aeolus-Mission**

# First year in space

European Space Agency

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# ESA-DEVELOPED EARTH OBSERVATION MISSIONS





Satellites

# (1) Objectives

#### **Scientific objectives**

- Improve the quality of weather forecasts
- Advance the understanding of atmospheric dynamics and climate processes

#### **Explorer objectives**

 Demonstrate space-based Doppler Wind LIDARs potential for operational use

#### **Observation means**

 Provide global measurements of horizontal line of sight (HLOS) wind profiles in the troposphere and lower stratosphere



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## (1) Forecast bust, Europe 03/14





## (2) Aeolus – Mission design

Polar orbit Ascending node 1800LT 7day repeat cycle, 111orbits

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Dins

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320km

- Sun

Aeolus

230km

\_35°



)km Aeolus

Europea 🔶

European Space Agency

Sun





## (2) Aeolus – ALADIN (Atmospheric Laser Doppler Instrument)





- > Nd:YAG
- > Diode pumped
- > Wavelength 354.8nm
- Repetition rate 50.5Hz
- Emit energy 80mJ

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## (2) Aeolus – ALADIN (Atmospheric Laser Doppler Instrument)





## (3) Launch and first measurement





## (3) Measurement example





https://www.esa.int/Our\_Activities/Observing\_ the\_Earth/Aeolus/Watch\_Aeolus\_launch\_replay



L2B Rayleigh-clear and Mie-cloudy results from file: p3/working/orbit 5947 5947 sappa/AE OPER ALD U N 2B 20190901T224235 20190902T001235 0001.TXT



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## (3) Fit of short-range forecasts to other obs.





## (3) Inflight Observations:



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## (3) Inflight Observations: Emit Energy





### **Energy decrease** FMA

- Master oscillator alignment
- Photodiode calibration factor
- Laser diode degradation
- $\rightarrow$  Can be revisited later

#### <u>FMB</u>

Under investigation

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## (3) Inflight Observations: Frequency Noise





### **Frequency noise**

- 1 m/s ≈ 5.64 MHz
- 5-7 MHz rms achieved in orbit!



Courtesy of A. Ciapponi, ESTEC

- (yet) uncorrelated time periods of increased frequency jitter (100MHz p-p)
- Orbital variation found
- 11-15 MHz rms nadir attitude

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## (3) Inflight Observations: Hot Pixels



#### **Hot Pixels**

- Continuously increasing number of pixels show increased dark current
- Longterm in the order of 2-15LSB  $\approx$  2e-/shot, partially RTS signature



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## (3) Inflight Observations: Hot Pixels



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### Summary



- ✓ More than 13 years of development <u>challenges</u>
- ✓ Invaluable <u>experience</u> has been gained, e.g.,
  - Laser thermal-mechanical design (energy, pointing drifts) and susceptibility of laser frequency to microvibrations
  - Optical component development and testing for high UV-energy lasers
  - Telescope thermal-mechanical design
  - Lidar system-level design (laser + optical receiver) and its representative end-to-end testing (OGSE as atmospheric simulator)





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### Summary



From a lidar-instrument perspective it was demonstrated that:

- a space-borne, direct-detection wind lidar can measure atmospheric winds by use of molecular Rayleigh and aerosol/cloud Mie backstatter => technical proof
- a powerful UV lidar can be operated in space over 12 months with high frequency stability
- internal and atmospheric calibrations can be used to characterize the instrument including returns from the non-moving ground => bias corrections
- Positive impact on NWP according to first preliminary results as reported by various met centers



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## Aeolus – Range Bin Strategy



- Current settings will be kept global until October (data set July Sep)
  - Northern Hemisphere + Tropics (+90 to -25deg latitude, stratospheric aerosol)
  - South Mid-Latitudes (-25 to -60deg)
  - South Pole (-60 / -90deg, PSC)



 Campaign boxes with specific settings added (Sep to October)
 (DLR-AVATARI around Iceland,

East Mediterranean Aerosol Range bin Settings)

- Plans to perform special 3-week measurement period for AMV (Oct) (need 250m RBS in specific altitudes)
- New global settings Nov-Feb based on first feedback and for Strateole2 (latitudinal belts 90-60-30deg)

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## Major development lessons learned: 1) Changes of thermal interfaces in vacuum





Combining highly dissipative pump units and a highly alignment sensitive MO (without isostatic mounting) is simply not good design and should be changed



Relative energy evolution for different mechanical configurations of the Aladin laser pump units showing energy degradation of 30% over a few hundred hours (n.b. this only happens in vacuum when the pump units are ON). A residual element of this probably caused the energy loss on FM-A.

#### **Lessons learned:**

- Where possible remove units with high dissipation away from alignment sensitive components
- Attempt to minimize the mechanical constraints on the interface and/or the support structure for the alignment sensitive elements by making their mounts iso-static

#### **Design solutions:**

- Use ATLID design (PU's moved away from optical bench) with additional pump unit with predemonstration that this improves the stability
- Improve the iso-staticity of the mounts

## Major development lessons learned : 2) LID: control your processes and test as you fly



Laser-induced damage events during the Aladin laser development

#### Lessons learned:

- □ If you operate in vacuum, test in vacuum
- □ Use fully densified coatings
- □ Use processes that inherently reduce the number of small defects on both substrates and coatings
- □ Ensure adequate physical-chemical analysis is undertaken to avoid potential problems and control processes



55.93u /Fe+ MC: 11: TC: 1.430e+004

Micrographs Small defects can lead to damage after 10's of Mshots and they can arise from many sources

Good substrate

## Major development lessons learned : 3) LID: you are only as good as your weakest optic



#### >350 individual optics and >30 >100 optics tested and >50 different manufacturing processes flight optics screened





Instrument level testing with FM-A

#### Lessons learned:

- Test the full area of the optic which contains the beam
- If damage precursors are <sup>a</sup> activated then perform test directly on the defect to ensure it will be stable
  Test for an adequate number of shots
  Ensure there is a margin on LIDT by > x2



Instrument level testing of the weakest optic in the emission path

## Major development lessons learned : 4) You need some oxygen to prevent LIC

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LIC: highly absorbing deposits formed by the interaction of the laser with organic outgassing on the surface of optics



Loss of 50% of the energy of the Aladin EM laser vacuum test after 6 hrs (not good for a 3yr mission)

#### **Lessons learned:**

- □ Classical contamination control methods do not stop LIC but still mitigate its effect
- □ Get rid of "bad" materials (silicones, aromatics,... as far as you can)
- □ High power lasers need an oxidising environment to operate in



Lux et al., 2018





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Upper optical bench

Lower optical bench

| Parameter                   | Value              |
|-----------------------------|--------------------|
| Dimensions                  | (582x422x215)mm    |
| /olume                      | 30L                |
| Mass                        | 30Kg               |
| Power consumption           | 300W               |
| Wavelength                  | 354.8nm            |
| Pulsewidth                  | 20ns (FWHM)        |
| Pulse repitition frequency  | 50.5Hz             |
| Energy (IR)                 | 280mJ              |
| Energy (UV)                 | 80mJ               |
| Harmonic section conversion |                    |
| efficiency                  | 28%                |
| Peak fluence (UV)           | 1Jcm <sup>-2</sup> |
| Wall plug efficiency (UV)   | 1%                 |

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## (1) Wind information in recent ECMWF cycle



log<sub>10</sub>(Number of u-component obs per 10<sup>6</sup>km<sup>2</sup>)



| Obs type          | % of<br>total<br>count | Mean assigned<br>u-wind error<br>(m/s) |
|-------------------|------------------------|--|
| AMVs              | 47                     | 4.6                                    |
| Scattero<br>meter | 23                     | 1.5                                    |
| Radiosondes       | 11                     | 2.0                                    |
| Wind<br>profilers | 10                     | 1.8                                    |
| Aircraft          | 9                      | 2.4                                    |



## (3) Inflight Observations: Harmonic Bias

### **Harmonic Bias**

 Orbital cycle of satellite to surface distance and thermo elastic variations result in harmonic residual of the "zero wind" surface return



- Seasonal variation
- Introduce bias up to ±6m/s that can be corrected to about 1m/s
  - Non-orbital biases under investigation

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